

Energy inequalities under one-sided curvature bounds

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joint work with

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Outline

1 Motivation

- Setting and Definitions
- Energy Estimates Using Two-Sided Bounds

2 Our Results

- Energy Estimates Using One-Sided Bounds
- Basic Ideas for Proof
- Further Results

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Curved Spacetime

- $(\mathcal{M}^{3+1}, \mathbf{g})$ is a globally hyperbolic spacetime, Ricci-flat
- foliated by compact or asymptotically flat hypersurfaces \mathcal{H}_t , time parameter $t \in I$
- \mathbf{N} future-directed time-like unit normal
- n lapse
- \mathbf{k} second fundamental form
- orthonormal frame $\{\mathbf{e}_0 = \mathbf{N}, \mathbf{e}_1, \mathbf{e}_2, \mathbf{e}_3\}$

Change of Metric in Normal Direction

Deformation tensor of the foliation

$$\pi = \mathcal{L}_{\mathbf{N}}\mathbf{g}$$

Components:

- $\pi_{00} = 0$
- $\pi_{0i} = \nabla_i \log n$
- $\pi_{ij} = -2k_{ij}$

Bel–Robinson Tensor

Bel–Robinson tensor

$$Q_{\alpha\beta\gamma\delta} = Q[\mathbf{R}]_{\alpha\beta\gamma\delta} := R_{\alpha\lambda\gamma\mu} R_{\beta}{}^{\lambda}{}_{\delta}{}^{\mu} + {}^*R_{\alpha\lambda\gamma\mu} {}^*R_{\beta}{}^{\lambda}{}_{\delta}{}^{\mu}$$

- can be decomposed using the electric and magnetic parts of the curvature tensor \mathbf{R} , e.g.

$$Q_{0000} = E_{ij}E^{ij} + H_{ij}H^{ij} = |\mathbf{E}|^2 + |\mathbf{H}|^2$$

$$E_{ij} = R_{i0j0} \quad H_{ij} = {}^*R_{i0j0}$$

- is divergence-free on Ricci-flat spacetimes: $\mathbf{D}^{\alpha} Q_{\alpha\beta\gamma\delta} = 0$

Bel–Robinson Energy

Total Bel–Robinson energy at time t

$$Q[\mathbf{R}]_{\mathcal{H}_t} := \int_{\mathcal{H}_t} Q[\mathbf{R}](\mathbf{N}, \mathbf{N}, \mathbf{N}, \mathbf{N}) dV_{\mathbf{g}_t}$$

WANT : control the Bel–Robinson energy at time t by its values on some initial slice

assuming a two-sided bound on the deformation tensor one obtains

Theorem (e.g. Wang 2012)

Suppose $\int_{t_0}^{t_1} \|\pi\|_{L^\infty(\mathcal{H}_s)} ds = K < \infty$ then there exists a constant $C = C(K, t_1) > 0$ such that for all $t \in [t_0, t_1)$

$$Q(t) \leq CQ(t_0).$$

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Main Result

Theorem (B., Grant, LeFloch 2012)

Given a vacuum spacetime endowed with a foliation $(\mathcal{H}_t)_{t \in I}$ with lapse n and second fundamental form \mathbf{k} , one has for $t_1 \leq t_2$, $t_1, t_2 \in I$,

$$Q[\mathbf{R}]_{\mathcal{H}_{t_2}} \leq e^{3K_{n,\mathbf{k}}(t_1,t_2)} Q[\mathbf{R}]_{\mathcal{H}_{t_1}},$$

where

$$K_{n,\mathbf{k}}(t_1, t_2) := \int_{t_1}^{t_2} \sup_{\mathcal{H}_t} \rho(n, \mathbf{k}) dt$$

and $\rho(n, \mathbf{k})$ is the largest eigenvalue of the symmetric 6×6 matrix $\Pi(n, \mathbf{k})$.

Sketch of Proof I

- express the Bel-Robinson tensor in terms of electric and magnetic parts of curvature to obtain

$$-\frac{1}{2}Q_{\alpha\beta 00}n\pi^{\alpha\beta} = \text{tr} \left(\begin{pmatrix} \mathbf{E} & \mathbf{H} \end{pmatrix} \Pi \begin{pmatrix} \mathbf{E} \\ \mathbf{H} \end{pmatrix} \right)$$

with a symmetric 6×6 matrix Π depending on n , \mathbf{k} and π

- the matrix Π has 3 real-valued double eigenvalues $a_1 \leq a_2 \leq a_3 =: \rho(n, \mathbf{k})$



$$-\frac{1}{2}Q_{\alpha\beta 00}n\pi^{\alpha\beta} \leq a_3 \text{tr} \left(\begin{pmatrix} \mathbf{E} & \mathbf{H} \end{pmatrix} \begin{pmatrix} \mathbf{E} \\ \mathbf{H} \end{pmatrix} \right) = \rho(n, \mathbf{k})Q_{0000}$$

Sketch of Proof II

- since the Bel–Robinson tensor is divergence-free we have $\mathbf{D}^\alpha Q_{\alpha 000} = \frac{3}{2} Q_{\alpha\beta 00} \pi^{\alpha\beta}$
- and Stokes' theorem implies on $\mathcal{M}_{[t_1, t_2]} = \bigcup_{t \in [t_1, t_2]} \mathcal{H}_t$:

$$\begin{aligned} Q_{\mathcal{H}_{t_1}} - Q_{\mathcal{H}_{t_2}} &= \frac{3}{2} \int_{\mathcal{M}_{[t_1, t_2]}} Q_{\alpha\beta 00} \pi^{\alpha\beta} dV_{\mathbf{g}} \\ &\geq -3 \int_{t_1}^{t_2} \sup_{\mathcal{H}_s} (\rho(n, \mathbf{k}) Q_{\mathcal{H}_s}) ds. \end{aligned}$$

- Gronwall argument: $Q_{\mathcal{H}_{t_2}} \leq e^{3 \int_{t_1}^{t_2} \sup_{\mathcal{H}_t} \rho(n, \mathbf{k}) dt} Q_{\mathcal{H}_{t_1}}$

Generalization to Weyl Fields

Weyl field

A Weyl field is a $(0, 4)$ -tensor field $W_{\alpha\beta\gamma\delta}$ that has the same symmetries as the Riemann tensor and is trace-free, i.e.

$$W_{\alpha\beta\gamma\delta} = W_{[\alpha\beta][\gamma\delta]} \quad W_{[\alpha\beta\gamma]\delta} = 0 \quad W_{\beta\alpha\delta}^{\alpha} = 0.$$

Bel-Robinson tensor **and** energy:

$$Q[\mathbf{W}]_{\alpha\beta\gamma\delta} := W_{\alpha\lambda\gamma\mu} W_{\beta}^{\lambda}{}_{\delta}{}^{\mu} + {}^*W_{\alpha\lambda\gamma\mu} {}^*W_{\beta\delta}{}^{\mu}$$

$$\mathcal{Q}[\mathbf{W}]_{\mathcal{H}_t} := \int_{\mathcal{H}_t} Q[\mathbf{W}](\mathbf{N}, \mathbf{N}, \mathbf{N}, \mathbf{N}) dV_{g_t}$$

Result on Weyl Fields

Theorem (B., Grant, LeFloch 2012)

Same notation as before. Any Weyl field \mathbf{W} defined on \mathcal{M}^{3+1} satisfies

$$\begin{aligned} Q[\mathbf{W}]_{\mathcal{H}_{t_2}} &\leq e^{3K_{n,k}(t_1, t_2)} Q[\mathbf{W}]_{\mathcal{H}_{t_1}} \\ &\quad - \int_{t_1}^{t_2} e^{3K_{n,k}(t, t_2)} \int_{\mathcal{H}_t} n(\operatorname{div} Q[\mathbf{W}])_{000} dV_{\mathbf{g}_t} \end{aligned}$$

for $t_1 \leq t_2$, $t_1, t_2 \in I$.

Summary

- The growth of the Bel–Robinson energy can be controlled by the lapse and the second fundamental form of the foliation.
- Algebraic properties of the Bel–Robinson tensor show that one-sided geometric bounds are sufficient.
- Other observations
 - similar results can be obtained using Maxwell and Yang–Mills Fields
 - bounds of the deformation tensor play also a role in the formulation of breakdown criteria for the vacuum Einstein equations (Klainerman–Rodnianski 2010, Wang 2012)